

Unruh effect and Quantum Computing

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discussing with

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Synthetic Unruh effect in cold atoms

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We propose to simulate a Dirac field near an event horizon using ultracold atoms in an optical lattice. Such a quantum simulator allows for the observation of the celebrated Unruh effect. Our proposal involves three stages: (1) preparation of the ground state of a massless 2D Dirac field in Minkowski spacetime; (2) quench of the optical lattice setup to simulate how an accelerated observer would view that state; (3) measurement of the local quantum fluctuation spectra by one-particle excitation spectroscopy in order to simulate a De Witt detector. According to Unruh's prediction, fluctuations measured in such a way must be thermal. Moreover, following Takagi's inversion theorem, they will obey the *Bose-Einstein* distribution, which will smoothly transform into the *Fermi-Dirac* as one of the dimensions of the lattice is reduced.

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The Unruh effect and its applications

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It has been thirty years since the discovery of the Unruh effect. It has played a crucial role in our understanding that the particle content of a field theory is observer dependent. This effect is important in its own right and as a way to understand the phenomenon of particle emission from black holes and cosmological horizons. Here, we review the Unruh effect with particular emphasis to its applications. We also comment on a number of recent developments and discuss some controversies. Effort is also made to clarify what seems to be common misconceptions.

Unruh effect

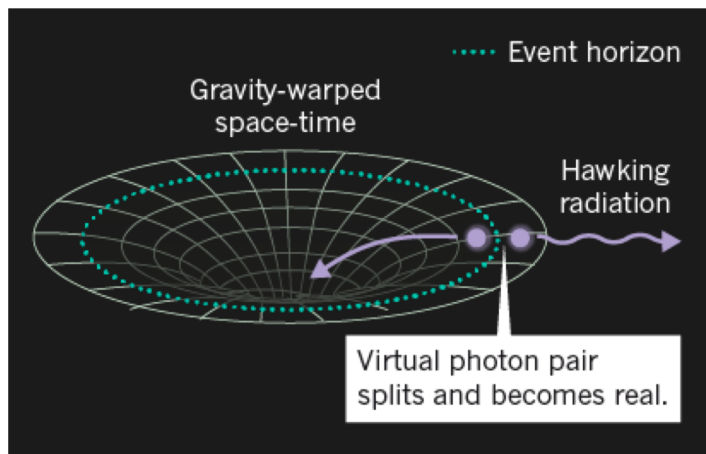
- An accelerated observer through the flat Minkowski spacetime vacuum will observe a thermal bath, a : constant acceleration

$$k_B T_U = \frac{\hbar a}{2\pi}.$$

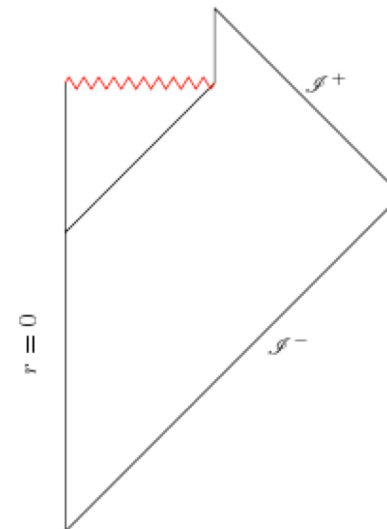
- Similar to Hawking radiation
- Event Horizon (Black wall vs Schwarzschild radius) prevents communication between different region of space time, which plays a role of heat bath
- Thermal effect without underlying stochasticity
→ Thermal mixed state from pure quantum state

Hawking radiation

- Black Hole of mass M
- Vacuum fluctuation, particle-antiparticle pair creation/annihilation, near Schwarzschild radius
- one of pair escape from BH, and the other fall into BH
- Emitted particle spectrum turns out to be blackbody radiation with temperature
$$T = \frac{\hbar c^3}{8\pi k G M}$$
- Blackhole lose mass (energy conservation), eventually evaporate, information paradox



[Nature]



[Wikipedia]

Rindler spacetime

- Flat Minkowski spacetime
- Consider an observer with constant acceleration $a=1$ in x-axis. Observer is at rest at $t=0$ and $x=1$
- Natural coordinate for the observer is Rindler coordinate (co-moving coordinate) $(t,x) \rightarrow (\xi, \eta)$

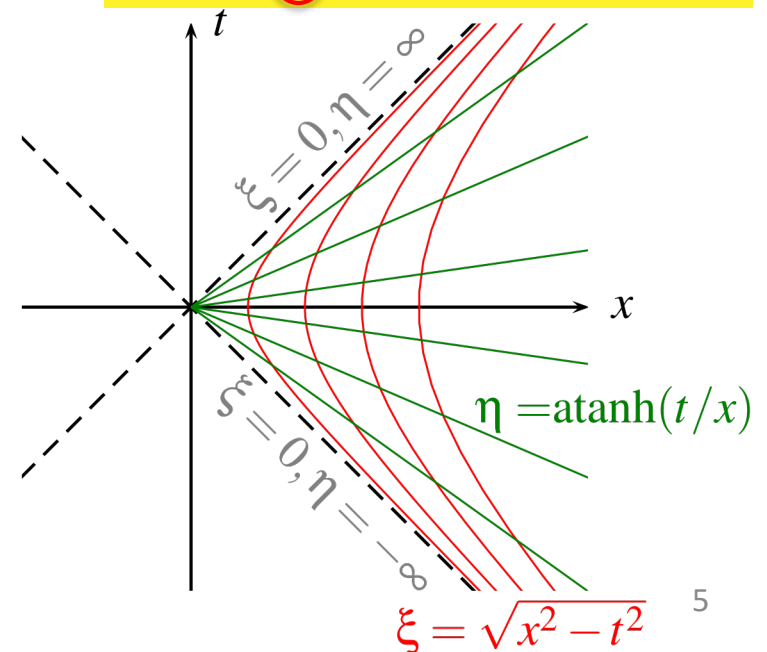
$$\begin{cases} t = \xi \sinh \eta \\ x = \xi \cosh \eta \end{cases}$$

- The observer's trajectory is $\xi=1$, and all η

- The Minkowski metric in Rindler coordinate

$$ds^2 = -\xi^2 d\eta^2 + d\xi^2 + dy^2 + dz^2$$

- spacetime is separate into two parts, left and right Rindler wedges, doesn't communicate each other
- From observer's point of view ($\xi=1$), light moves at $\xi < 1$ is slower than 1.
- At $\xi=0$, light stops (Black wall).



Thermalization theorem

- Spacetime with time-like Killing vector (metric is invariant along with the vector) has well defined energy
- Thermalization theorem (Tolman-Ehrenfest):

For any thermal equilibrium field, local temperature times time-like Killing vector is constant:

$$T \cdot g_{00}^{1/2} = \text{const}$$

c.f. redshift frequency

$$\nu(P') = \nu(P) \cdot \sqrt{g_{00}(P)/g_{00}(P')}$$

- Vacuum of Massless Dirac field (Fermi sea)

$$|0_M\rangle = \prod_{\omega_k^M < 0} b_k^\dagger |\Omega\rangle$$

- Vacuum of accelerated observable, Hamiltonian H_R

$$|0_R\rangle = \prod_{\omega_q^R < 0} d_q^\dagger |\Omega\rangle$$

- Minkowski vacuum is not vacuum (either eigenstate) of H_R
- For the observer, field in left Rindler Wedge (LRW) is not causal, state of LRW dof is traced out (loss of information)

$$\rho_R = \text{Tr}_L |0_M\rangle \langle 0_M|$$

- Density matrix can always formally written as a thermal state using entanglement Hamiltonian H_E $\rho_R = \exp(-H_E)$

- H_E and H_R commute according to boost invariance of Minkowski vacuum

$$0 = \dot{\rho}_R = -\frac{i}{\hbar} [\rho_R, H_R]$$

- It can be explicitly proved H_E is proportional to H_R

$$\rho_R = \exp\left(-\frac{H_R}{k_B T_U}\right) \quad k_B T_U = \frac{\hbar a}{2\pi}.$$

- Minkowski vacuum turns out stationary to an accelerated observer

Unruh effect

- Two different mode expansions of the same fermion field with definite wavefunctions
- Bogoliubov transformation $d_q^\dagger = \sum_k U_{qk} b_k^\dagger$
- Occupation of Rindler mode on the ^kMinkowski vacuum (IR cutoff by radius ϵ)

$$n_{q,\mathbf{r}} \equiv \int_{D_{\mathbf{r},\epsilon}} \langle 0_M | d_q^\dagger | \mathbf{r}' \rangle \langle \mathbf{r}' | d_q | 0_M \rangle = \sum_{\omega_k^M < 0} |\tilde{U}_{qk}(\mathbf{r})|^2,$$

$$n_{q,\mathbf{r}} = \frac{1}{\exp(\hbar\omega_q^R/k_B T_U(\mathbf{r})) + 1}$$

$$k_B T_U(\mathbf{r}) = \frac{\hbar}{2\pi x}$$

- De Witt detector (minimally coupled to fermion field at \mathbf{r}), whose absorption/emission is written by Wightman function

$$G(t) \equiv \langle 0_M | c_{x(t)}^\dagger(t) c_{x(0)}(0) | 0_M \rangle$$

$x(t)$: trajectory of the observer, c_x is the annihilation operator at x

- detector response function $G(\omega)$, Fourier transformation of $G(t)$

Unruh effect (contd.)

- Minkowski operator b_k in terms of coordinate space operator c_x
(M_{kx} : wave function = plane wave)

$$b_k^\dagger = \sum_x M_{kx} c_x^\dagger$$

- Rindler operator d_q in terms of coordinate space operator c_x
(R_{qx} : wave functions)

$$d_q^\dagger = \sum_x R_{qx} c_x^\dagger$$

- c.f. Bogoliubov transformation

$$d_q^\dagger = \sum_k U_{qk} b_k^\dagger \quad U_{qk} = \sum_x R_{qx} \bar{M}_{kx}$$

- Detector response function

$$G_{x_0}(\omega) \equiv \int dt e^{-i\omega t} \langle 0_M | c_{x_0}^\dagger(t) c_{x_0}(0) | 0_M \rangle$$

$$= \sum_{q,q'} \delta(\omega - \omega_q^R) \bar{R}_{qx_0} R_{q'x_0} \sum_{\omega_k^M < 0} \bar{U}_{qk} U_{q'k} = \frac{1}{\exp(\hbar\omega_q^R/k_B T_U(\mathbf{r})) + 1} \quad k_B T_U(\mathbf{r}) = \frac{\hbar}{2\pi x}$$

is the Fermi-Dirac thermal distribution (in 1+1 dim and 3+1 dim)

is **Bose-Einstein** distribution (in 2+1 dim) for Fermion, and opposite for Boson (Takagi's inverse theorem)

$$G(\omega) = \frac{1}{\exp(\omega/T(x)) - 1}, \quad T(x) = \frac{1}{2\pi x}$$

Dirac fermion in Rindler Lattice

- Minkowski Dirac Hamiltonian

$$i\partial_0\psi = \mathcal{H}\psi = -i\gamma_0\gamma^j\partial_j\psi \quad i\partial_t\psi = -i(\partial_x\sigma_x + \partial_y\sigma_y)\psi$$

- curved spacetime Hamiltonian (w_μ : spin connection)

$$\partial_\mu\psi \rightarrow D_\mu\psi \equiv \left(\partial_\mu + \frac{1}{4}w_\mu^{ab}\gamma_{ab}\right)\psi \quad \gamma_{ab} \equiv \frac{1}{2}[\gamma_a, \gamma_b].$$

$$i\partial_t\psi = -i\gamma_t\left(\gamma^j\partial_j + \frac{1}{4}\gamma^jw_j^{ab}\gamma_{ab} + \frac{1}{4}\gamma^tw_t^{ab}\gamma_{ab}\right)\psi$$

- Rindler coordinate $ds^2 = -x^2dt^2 + dx^2 + dy^2$

- spin connection (others are zero) $w_t^{01} = x/|x|$

- Dirac equation and Hamiltonian density

$$i\partial_t\psi = -i\left(\left(|x|\partial_x + \frac{1}{2}\frac{x}{|x|}\right)\sigma_x + |x|\partial_y\sigma_y\right)\psi \quad \mathcal{H}_R = -i\left(\left(|x|\partial_x + \frac{1}{2}\frac{x}{|x|}\right)\sigma_x + |x|\partial_y\sigma_y\right)$$

$$\mathcal{H}_R = \sqrt{x}\not{p}\sqrt{x}.$$

- Second quantized Hamiltonian $H_R = \int dx dy \bar{\psi}^\dagger \mathcal{H}_R \psi$

- $x=0$ is the **boundary** (any boundary condition is allowed at $|x|=0$)

Eigenstate of Rindler Hamiltonian in 1+1 dim

- 1-particle Hamiltonian (density) $H_{R(1D)} = \sqrt{x} p \sqrt{x}$
- variable transformation $u = \log(x)$. (boundary $x=0$ is $u \rightarrow -\infty$)

$$H_{R(1D)} = -i(\partial_u + 1/2)$$

- eigenvalue equation

$$-i(x\partial_x + 1/2)\psi(x) = -i(\partial_u + 1/2)\psi(u) = \omega\psi(u)$$

- eigen functions (plane waves in $u = \log(x)$)

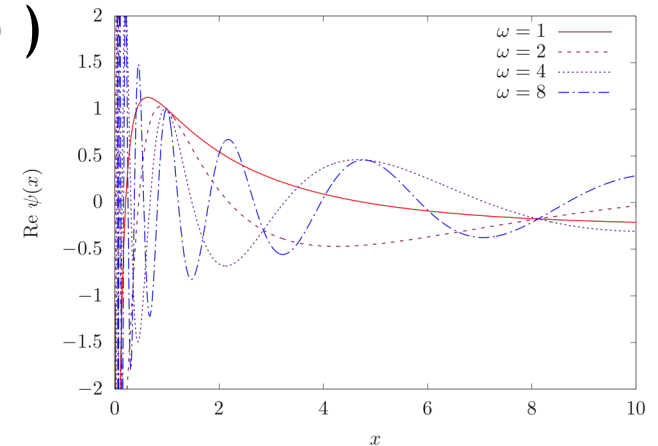
$$\psi(u) = A \exp \left[\left(i\omega - \frac{1}{2} \right) u \right] = A x^{i\omega - 1/2}$$

orthonormal

$$\int_0^\infty dx \exp((-i\omega - 1/2)u) \exp((i\omega' - 1/2)u) = \int_{-\infty}^\infty du \exp(-i\omega u) \exp(i\omega' u) = \delta(\omega - \omega').$$

- Spinor structure

$$-i(x\partial_x + 1/2)\sigma_x \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix} = \omega \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix} \quad \begin{pmatrix} \psi_1(x, t) \\ \psi_2(x, t) \end{pmatrix} = A \begin{pmatrix} 1 \\ \pm 1 \end{pmatrix} x^{i\omega - 1/2} e^{-i\omega t}$$



Lattice Discretization

- 1+1 dim $H_{R(1D)} = \sqrt{x} p \sqrt{x}$
- $x_m = m d$, d : lattice spacing, $m = -(L-1)/2, \dots, (L-1)/2$, even L
- symmetric derivative for $p = -i\partial_x$

$$(p\psi)_m = -i (\psi_{m+1} - \psi_{m-1}) / (2d)$$

- 1 particle Hamiltonian $H_{R(1D)} = \sqrt{x} p \sqrt{x}$

$$R_{m,m+1} = -\frac{i}{2} \sqrt{m(m+1)}$$

- for 2+1 dim

$$H_R = -\sum_{m,n} t'_0 \left(\left(m + \frac{1}{2}\right) e^{i\frac{\pi}{2}(m-n)} c_{m+1,n}^\dagger + m e^{i\frac{\pi}{2}(m-n)} c_{m,n+1}^\dagger \right) c_{m,n} + \text{H.c.},$$

Simulating Unruh effect

- Detector response function $G(\omega)$, Fourier transformation of

$$G(t) \equiv \langle 0_M | c_{x(t)}^\dagger(t) c_{x(0)}(0) | 0_M \rangle$$

- Measure overlap between the states of one-hole excitation at $t=0$ and t
- In Right Rindler coordinate, observer's trajectory is $x(0) = x(t) = x_0$ (frame independence, equivalence principle)
- Strategy : quench from Minkowski Hamiltonian to Rindler spacetime
 1. Prepare Minkowski vacuum
 2. Quench : Switch Hamiltonian from Minkowski H_M to Rindler H_R
quench introduces event horizon, disconnecting the L/R Wedges
 3. Measure $G(t)$ in Rindler frame at $x=x_0$

Numerical simulation

- 2+1 dim and 1+1 dim
- space size L_x, L_y
- time step t_0 for Minkowski
- time step t_0' for Rindler
chose $t_0' = 2 t_0 / L_x$
- Lattice spacing $d = t_0$

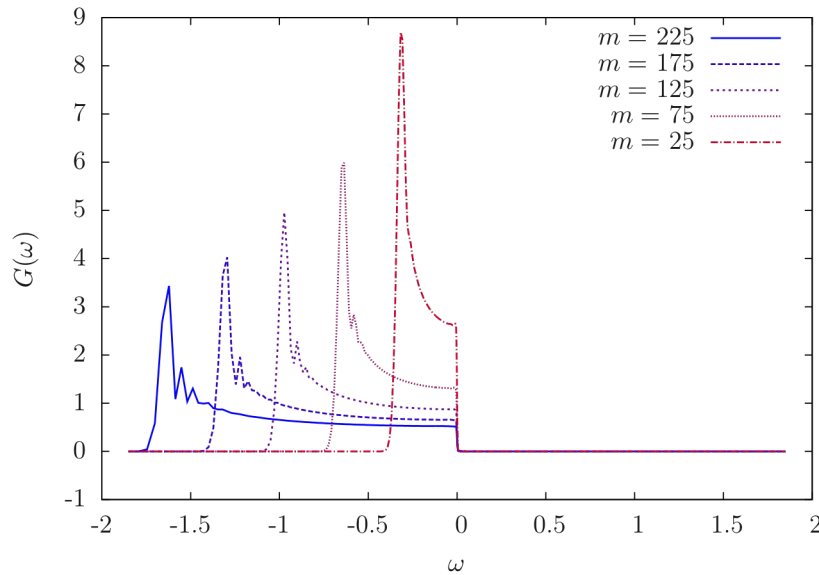
$$H_M = - \sum_{m,n} t_0 \left(e^{i \frac{\pi}{2} (m-n)} c_{m+1,n}^\dagger + e^{i \frac{\pi}{2} (m-n)} c_{m,n+1}^\dagger \right) c_{m,n} + \text{H.c.}$$

$$H_R = - \sum_{m,n} t_0' \left(\left(m + \frac{1}{2}\right) e^{i \frac{\pi}{2} (m-n)} c_{m+1,n}^\dagger + m e^{i \frac{\pi}{2} (m-n)} c_{m,n+1}^\dagger \right) c_{m,n} + \text{H.c.},$$

expectation $G(\omega) = \frac{1}{\exp(\omega/T(x)) + 1}, \quad T(x) = \frac{1}{2\pi x}$

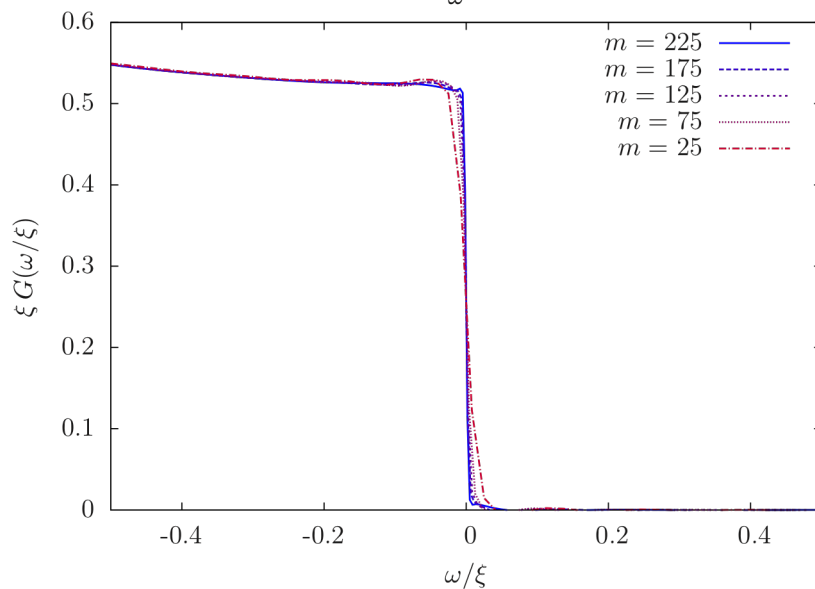
$$\xi G(\omega/\xi) = \xi \frac{1}{\exp(\pi L_x \omega) + 1}, \quad \xi = 2x/L_x$$

- $L_x = 500$, 1+1 dim
- $x = md$ (smaller x is hotter temperature)
- $|\omega / t_0| > 1$ is time-step discretization

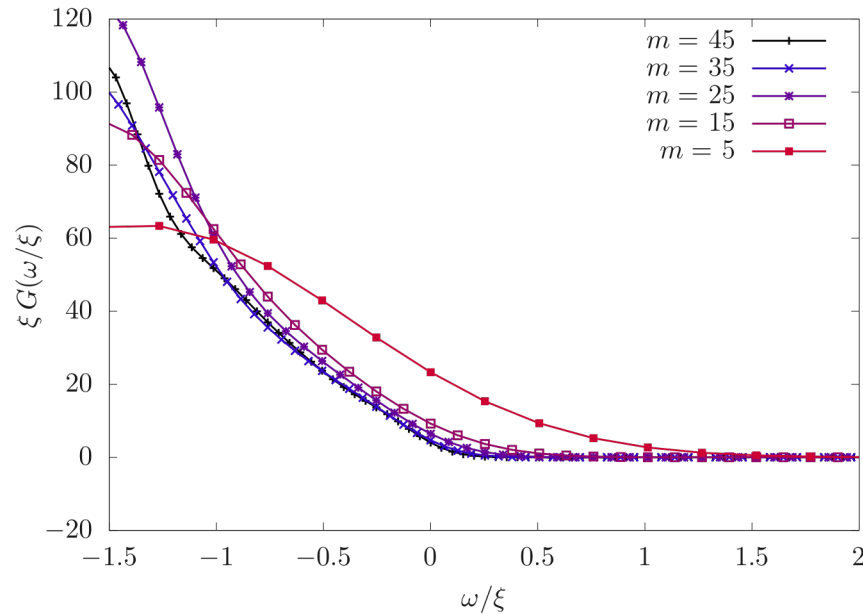


$$G(\omega) = \frac{1}{\exp(\omega/T(x)) + 1}, \quad T(x) = \frac{1}{2\pi x}$$

$$\xi G(\omega/\xi) = \xi \frac{1}{\exp(\pi L_x \omega) + 1}, \quad \xi = 2x/L_x$$



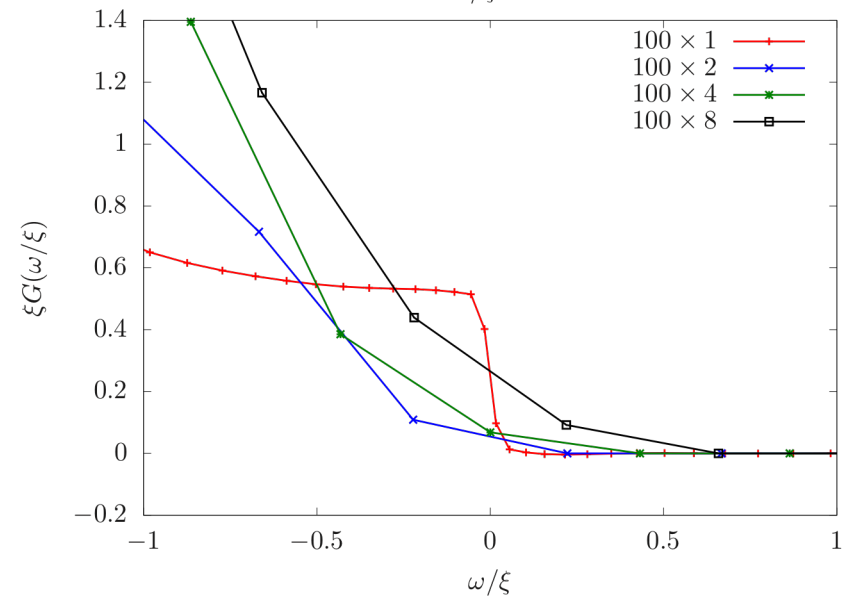
- 2+1 dim $L_x = L_y = 100$
- Bose Einstein distribution



- transition from 1+1 dim to 2+1 dim

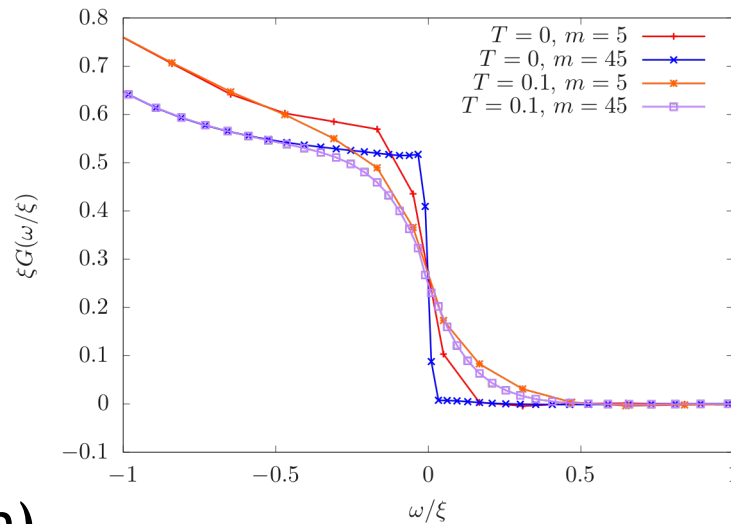
$$G(\omega) = \frac{1}{\exp(\omega/T(x)) - 1}, \quad T(x) = \frac{1}{2\pi x}$$

$$\xi G(\omega/\xi) = \xi \frac{1}{\exp(\pi L_x \omega) - 1}, \quad \xi = 2x/L_x$$

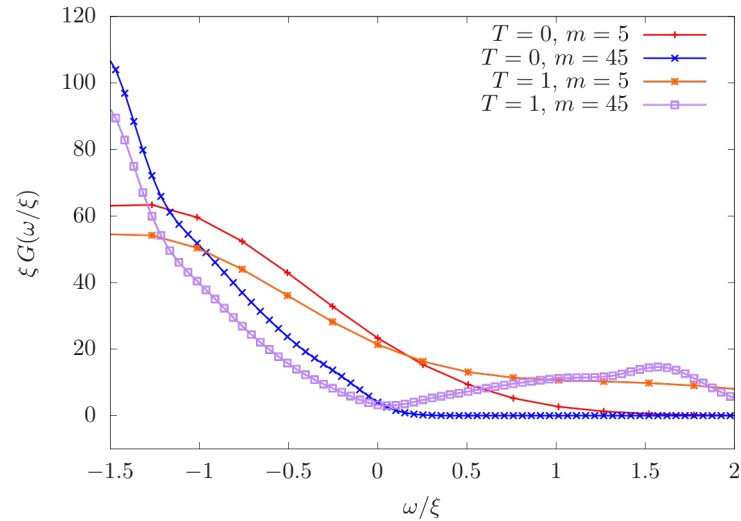


Discretization error

- 1+1 dimension (Fermi-Dirac)

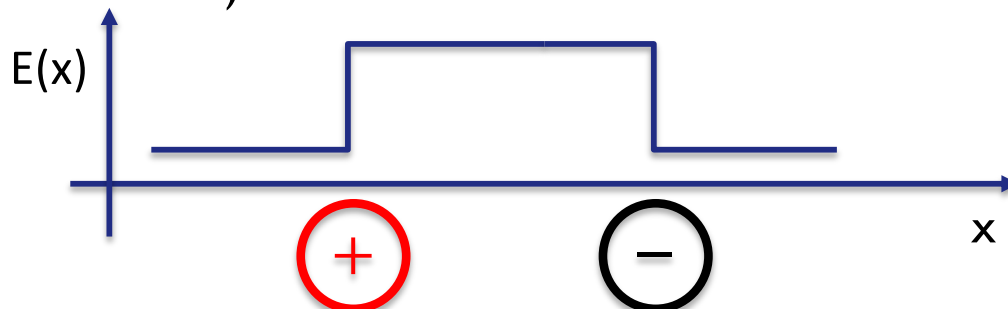


- 2+1 dim (Bose-Einstein)



Discussions and Prospects

- Authors proposed optical lattice experiment
- Apply this in digital quantum computation?
Number of Qbits too demanding ?
- Unruh effect for interacting theory such as Schwinger model?
(radiation for photon and electron ?)



- Pure quantum state, namely Minkowski vacuum, turns into thermal mixed state preparation, which is exponentially large operation otherwise.
- Minkowski vacuum could be prepared via adiabatic deformation from solvable Hamiltonian (Gumaro, Yuuta et al are now trying ultra-efficient preparation)
- Could we engineer different coordinate system so that the temperature is constant everywhere ?